Brief GlueX Note Original: January 9, 2008 Updated: January 9, 2008

## Reconciling Photoelectron Yield Estimates for the BCAL Test Module Alex R. Dzierba

This note summarizes some of the information we have on photoelectron yields  $(N_{pe})$  from measurements with cosmic rays and photon beam. We also examine to what extent these numbers are consistent.

## $N_{pe}$ from photon data

Alex Estimate - method 1: I estimated  $N_{pe}$  per end of the prototype BCAL module from the beam test Run 2334 - beam normally incident at the center. For each of eight bins in beam energy, from 200 to 600 MeV, I plotted the distribution in the ratio of the north sum to the south sum and fitted the resulting distribution, for each energy bin, to a Gaussian, yielding  $\mu_R$  and  $\sigma_R$ . Under the assumption that the number of photoelectrons per end,  $N_{pe}$  is given by:

$$N_{pe} = 2\frac{\mu_r^2}{\sigma_R^2}$$

the photoelectron yield per end is plotted in Figure 1 as a function of beam energy. A fit to a straight line:

$$N_{pe} = a + b \cdot E(MeV)$$

yields  $a = 14 \pm 4$  and  $b = 0.634 \pm 0.01$ . Extrapolating to 1 GeV gives us  $647 \pm 10$  photoelectrons.

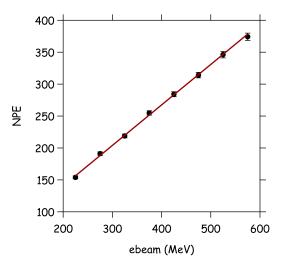


Figure 1: The number of photoelectrons per end of the BCAL module as a function of energy.

Alex Estimate - method 2: For the highest energy point (575 MeV) in Figure 1, the number of photoelectrons is  $374 \pm 6$ . Another approach is to apply the technique to each of the 18 channels to get a number of photoelectrons per channel and then add. The total number of photoelectrons in the 6 channels in the center of the module (channels 7 through 12) is  $373 \pm 7$ . Summed over all channels the total number of photoelectrons is  $440 \pm 8$ . Applying the scaling factor of 440/374 = 1.18 to the  $647 \pm 10$  photoelectrons at 1 GeV found above, the new number is  $761 \pm 12$  photoelectrons at 1 GeV.

**Blake Estimate:** Blake did a similar calculation (e-mail communication on 1/7/2007) using a Poisson fit rather than a Gaussian fit. He finds that at 500 MeV  $N_{pe} = 292$ . At that energy Alex's estimate is  $N_{pe} = 330 \pm 8$ .

**Comparison with KLOE:** Alex and Blake get consistent  $N_{pe}$  estimates. Alex's extrapolations to 1 GeV lie between 650 and 760 photoelectrons. KLOE quotes  $N_{pe} \sim 700$ .

# $N_{pe}$ from cosmic ray data:

Andrei estimates  $N_{pe} \sim 25$  per end per channel for events where the cosmic ray traverses three channels, one on top of the other. I note here that one of the KLOE studies of blue fibers coupled directly to a bi-alkali photocathode 2 m from a crossing minimum ionizing particle yields about 2 photoelectrons per mm of traversed fiber (actually they quote 1.7 to 2.25. Using this, one would expect  $N_{pe} \sim 38$  per channel in our case compared to our observed  $N_{pe} \sim 25$ . The difference may be due to coupling of the fiber to phototube.

#### Reconciling photon data and cosmic ray data:

In an earlier analysis Alex finds that for contained cosmic rays in a single layer (a layer consists of three channels - one on top of the other) the summed ADC distribution peaks at 1007 channels. From beam data, Alex and Blake find a consistent overall conversion of ADC counts to energy - Alex's number is 0.107, that is, energy =  $0.107 \times$  ADC counts. So a vertical penetrating cosmic ray is equivalent to a photon horizontally incident with energy 107 MeV. Based on Andrei's estimate of  $N_{pe} \sim 75$  for three channels, a 107 MeV photon should yield  $N_{pe} \sim 75$  per end. Extrapolating this to 1 GeV corresponds to ~ 700 photoelectrons.

A vertically incident cosmic ray passes through 3 channels. Each channel is 3.8 cm thick and on average corresponds to 1.9 cm of scintillating fiber. Assuming an energy loss of 2 MeV/cm in scintillator, that corresponds to 3.8 MeV per channel or 11.4 MeV for three channels. The equivalence of a vertically passing cosmic ray to a 107 MeV photon corresponds to a sampling fraction of about 11%, assuming more or less full shower containment for the photon.

### Photoelectron yield estimates for cosmic data:

Based on KLOE numbers we should expect  $N_{pe} \sim 38$  per channel for cosmic rays. Can we understand this from first principles? For scintillator, the photon yield per MeV of deposited energy is  $N_0 = 8000$  photons/MeV. We showed above that the energy deposited in scintillator per channel is  $E_{dep} = 3.8$  MeV. From the Bicron brochure, the minimum capture ratio for single-clad fibers is  $f_{cap} = 0.034$ . In our study of the emission spectra for blue and green fibers (GlueX-doc-924) we saw that for a XP2020 coupled to blue fiber 2 m from the source, the fraction of light surviving is  $f_{surv} = 0.24$ . In that same note we estimate the average XP2020 quantum efficiency to be QE = 0.14 assuming a blue fiber. So the  $N_{pe}$  should be:

$$N_{pe} = N_0 \cdot E_{dep} \cdot f_{cap} \cdot f_{surv} \cdot QE = 35$$

For BCAL we are using double-clad as opposed to single clad fibers and the improvement in the capture ratio should be 5/3 which would yield  $N_{pe} \sim 58$ . The light guide used for the BCAL test may have a transmission of ~ 0.6 which would get us back to  $N_{pe} \sim 38$ .